# Effect of Incompressibility and Symmetry Energy Density on Charge Distribution and Radii of Closed-Shell Nuclei

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#### Abstract

In this work, the effect of incompressibility modulus  $K_{NM}$  and symmetry energy density J on charge distribution and root-mean-square radii of neutron  $R_n$  and proton  $R_p$  has been investigated for light, medium and heavy closed-shell nuclei  $\frac{40}{20}Ca_{20}$ ,  $\frac{48}{20}Ca_{28}$ ,  $\frac{90}{40}Zr_{50}$ ,  $\frac{116}{50}Sn_{66}$ ,  $\frac{124}{62}Sm_{82}$  and  $\frac{208}{62}Pb_{126}$  within the framework of self-consistent Hartree-Fock (HF) with 20 types of Skyrme interactions having wide range of nuclear properties such as incompressibility modulus  $K_{NM}$  and symmetry energy density J. The nuclear charge densities have been obtained and compared with the experiment data to give us a picture about the internal structure of the investigated nuclei. Also, the relation between the proton and neutron root-mean-square radii, and nuclear matter properties has been studied and examined by determining and discussion the statistical Pearson linear correlation coefficient, to help in formulating nuclear interaction.

# 1. Introduction:

The main problem in physics is many-body problem [1], In nuclear physics, the many-body systems and the nature of the nuclear force [2] are the most important challenges facing physics the theoretical nuclear physics, therefore, theoretical researches in nuclear physics aim to describe the structure of the nuclei with a universal nuclear model.

The properties of ground-state of atomic nuclei can be studied by Hartree-Fock (HF) [3], [4] or Hartree-Fock- Bogoliubov [5], [6]. There are many models used to study the excited levels, and the Random Phase Approximation (RPA) is considered successful in studying the closed-shell nuclei [7]. Charge density distribution and radii are fundamental nuclear properties applied to probe the nucleus structure and describes the effect of effective interactions on nuclear structure [8], [9] which were measured by electron elastic scattering and muonic X-ray data [10] by methods based on the electromagnetic interaction between the nucleus and electrons or muons.

It has also been indicated by many research groups that

the nuclear properties are closely related to the symmetry term of the equation of state (EOS) [11], [12]. Charge radii is a good indicator of nuclear properties and the nuclear symmetry energy that characterizes the equation of state of matter [13], also, it found correlates strongly with other observables in nuclei, such as nuclear charge and radii [14],[15],[16], or neutron skin thickness [17], [18], [19], and their relationship to the nuclear matter properties [20], [21]. The present work aims to study the effect of incompressibility modulus  $K_{NM}$ and symmetry energy density J on charge distribution and root-mean-square radii of neutron  $(R_n)$  and proton  $(R_p)$  for light, medium and heavy closed-shell nuclei  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ , <sup>116</sup>Sn, <sup>144</sup>Sm and <sup>208</sup>Pb using the self-consistent Hartree-Fock (HF) with 20 types of Skyrme interactions. Having a large number of Skyrme-force parameterizations requires continuous search for the best to describe the experimental data, therefore, the nuclear charge densities to be obtained and compared with the experiment data to give us an imagination about the internal structure of the investigated nuclei, as well as, the dependence of proton and neutron root-mean-square radii to nuclear matter properties such as incompressibility modulus  $K_{NM}$  and symmetry energy density J to be examined by determining the statistical Pearson linear correlation coefficient.

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## 2. Theoretical framework:

The second quantization Hamiltonian of a many-body fermion system is a summation of the kinetic energy t is and the antisymmetrized two body matrix-elements

 $\bar{v}_{n\acute{n}m\acute{m}} = (k_n k_{\acute{n}} | V | k_m k_{\acute{m}} - k_{\acute{m}} k_m)$  as given below [22],

$$H = \sum_{n\acute{n}} t_{n\acute{n}} \hat{c}^{\dagger}_{\acute{n}} \hat{c}_{\acute{n}} + \frac{1}{4} \sum_{n\acute{n}m\acute{m}} \bar{v}_{n\acute{n}m\acute{n}} \hat{c}^{\dagger}_{\acute{n}} \hat{c}^{\dagger}_{\acute{n}} \hat{c}^{\dagger}_{\acute{n}} \hat{c}_{\acute{m}} \hat{c}_{\acute{m}}$$
(1)

Where  $\hat{c}$  and  $\hat{c}^{\dagger}$  are annihilation and creation operators, respectively with  $\hat{c}|\rangle = 0$ . The indices:  $n, \acute{n}, m, \acute{m}$  covers all degrees of freedom of all available single particle states.

The Hartree-Fock energy can be expressed as,

$$E^{HF} = \sum_{nn'} t_{nn'} \rho_{n'n} + \frac{1}{2} \sum_{nn'mm'} \rho_{nm} V_{nn'mm'} \rho_{m'n'}$$
(2)

where single-particle density  $\rho$  is diagonal which has the eigenvalues 1 or 0 for occupied and non-occupied states, and the energy should be minimized to determine the HF-basis [22],

$$h_{nn'}^{HF} = t_{nn'} + \Gamma_{nn'}(\rho) \tag{3}$$

with the self-consistent field.

$$\Gamma_{nn'}(\rho) = \sum_{mm'} V_{nn'mm'} \rho_{n'm'} \tag{4}$$

The Skyrme effective interaction [23], [24] is one of the most convenient forces used in HF calculations to describe the nuclear properties in ground-state, commonly contains central, non-local and density-dependent terms [25], [26]:

$$V(r_{1}, r_{2}) = t_{0}(1 + x_{0}P_{\sigma})\delta(r_{1} - r_{2})$$

$$+ \frac{1}{2}t_{1}(1 + x_{1}P_{12}^{\alpha}) \times [\vec{k}_{12}^{2}\delta(r_{1} - r_{2}) + \delta(r_{1} - r_{2})\vec{k}_{12}]$$

$$+ t_{2}(1 + x_{2}P_{12}^{\alpha})\vec{k}_{12}\delta(r_{1} - r_{2})\vec{k}_{12}$$

$$+ \frac{1}{6}t_{3}(1 + x_{3}P_{12}^{\alpha})\rho^{\alpha}(R)\delta(r_{1} - r_{2})$$

$$+ iW_{0}\vec{k}_{12}\delta(r_{1} - r_{2})(\vec{\sigma}_{1} + \vec{\sigma}_{2})\vec{k}_{12}$$
(5)

where

**.** . . /

 $\vec{k}_{12} = -\frac{i(\vec{\nabla}_1 - \vec{\nabla}_2)}{2}, \vec{k}_{12} = -\frac{i(\vec{\nabla}_1 - \vec{\nabla}_2)}{2}, P_{12}^{\alpha}$  is the spin-exchange operator and  $\sigma_i$  is the Pauli spin operator.

The Skyrme energy-density functional H(r) contains many terms and given by [27],

$$H = K + H_0 + H_3 + H_{eff} + H_{fin} + H_{so} + H_{sg} + H_{coul}$$
(6)

Where K is a kinetic-energy term,  $H_0$  is a zero-range term,  $H_3$  is a density-dependent term,  $H_{eff}$  is an effective-mass term,  $H_{fin}$  is a finite-range term,  $H_{sois}$  a spin-orbit term,  $H_{sg}$  is a tensor coupling with spin and gradient term and  $H_{coul}$  is a Coulomb interaction term. The charge density distribution in ground-state (in terms of Skyrme HF radial wave function u of the state nlj and occupation probability ) can be obtained by [28], [29].

$$\rho_q(r) = \frac{1}{4\pi} \sum_{nlj} \eta_q (2J+1) \left| \frac{u(nlj)}{r} \right|^2 \tag{7}$$

The root-mean-square (rms) radii is defined as,

$$r_q = \sqrt{\langle r_q^2 \rangle} = \left[ \int r^2 \rho_q(r) dr \middle/ \int \rho_q(r) dr \right]^{1/2}$$
(8)

$$\langle r_{ch}^2 \rangle = \langle r_p^2 \rangle + \langle r \rangle_p^2; \quad rms \quad radius \langle r \rangle_p = 0.8 fm \quad (9)$$

## 3. Result and discussion:

In this work, The HF calculations have been investigated for  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei by solving the equations numerically with 20 Skyrme-type interaction: SkM\* [30], SQMC650 [31], SV-K218 [32], SQMC700 [31], SkO [33], NRAPR [34], [35], T31 [36], Skχm [37], v075 [38], SkSC10 [39], RATP [40], SAMi [41], SK255 [42], SkD [43], GS2 [44], SIV [45], QMC2 [46], SIII [44], SVII [47]) the meaning of the abbreviation for each type of interaction and its parameters can be found in the cited references). Nuclear matter properties of the used Skyrme interaction such as the incompressibility modulus  $K_{NM}$  (MeV) and the symmetry energy density J (MeV), are presented in Table 1. Our results are consistent with those of HFB calculations [5],[6].

The nuclear charge densities have been obtained and compared with the experiment data to give us a picture about the internal structure of the investigated. Figs.(1, 2, 3 and 4) show the calculated HF charge distribution for  ${}^{40}Ca$  and  ${}^{48}Ca$ . Good agreements were obtained between our HF results for all of the used types of Skyrme interaction with the experimental data [8], [49] at the surface and interior regions, while our results underestimated the experimental data line inside the investigated nuclei SVII, SIII, QMC2 and GS2 sets.

For  ${}^{90}Zr$ , our results of charge distribution for all interactions were in a good agreement with the experimental data [8],[50] in the center and at the surface of the nucleus but not so good inside the nucleus as GS2, QMC2 and SVII interactions underestimated the experimental data line inside the nucleus and SQMC650, SQMC700, NRAPR, Skxm, SK255, T31, SAMi, RATP parameters in less overestimated the experimental data line inside the nucleus. There are many factors that control the success of the theoretical model, in Table 1, we notice that the value of  $K_{NM}$  for the first group is higher than that of the second group.

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Туре	Incompressibility	Symmetry
	Modulus	Energy Density
	$K_{NM}(MeV)$	J(MeV)
SkM*	216.610	30.03
SQMC650	218.110	33.65
SV-K218	218.230	30.00
SQMC700	222.200	33.47
SkO	223.340	31.97
NRAPR	225.700	32.78
KDE0	228.820	33.00
T31	230.010	32.00
Skχm	230.400	30.94
v075	231.290	28.00
SkSC10	235.890	22.83
RATP	239.520	29.26
SAMi	245.000	28.00
SK255	254.930	37.40
SkD	267.154	30.81
GS2	300.140	25.96
SIV	324.550	31.22
QMC2	330.100	28.70
SIII	355.370	28.16
SVII	366.440	26.96

**Table 1.** Nuclear matter properties of the used Skyrme interactions [48].

For <sup>116</sup>*Sn*, our results of charge distribution for all interactions were in a good agreement with the experimental data [8] in the center and at the surface of the nucleus but not so good inside the nucleus as in SQMC650, SKO,  $Sk\chi$ m parameters in less overestimated the experimental data line inside the nucleus and GS2, QMC2, SIII and SVII interactions underestimated the experimental data line inside the nucleus. The differences in describing data may due to differences in nuclear properties such as  $K_{NM}$ , where its value the first group is lower than of the second group.

For  $^{144}Sm$  our results of charge distribution for all interactions were in a good agreement with the experimental data [8] at the surface of the nucleus but not so good inside the nucleus.

For  $^{208}Pb$  our results of charge distribution for all interactions were in a good agreement with the experimental data [8] in the center and at the surface of the nucleus but not so good inside the nucleus as GS2, QMC2 and SVII and less for SIV and SIII interactions underestimated the experimental data line inside the nucleus and SQMC650, KDE0 parameters in less overestimated the experimental data line inside the nucleus. This may be due to differences in the nuclear matter properties of the investigated types of Skyrme interaction. By considering a light, medium and heavy nuclei, we can illustrate the impact of the effect of incompressibility modulus  $K_{NM}$  and symmetry energy density J on root-mean-square radii of neutron  $R_n$  and proton  $R_p$  for light, medium and heavy closed-shell nuclei  ${}^{40,48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$ .

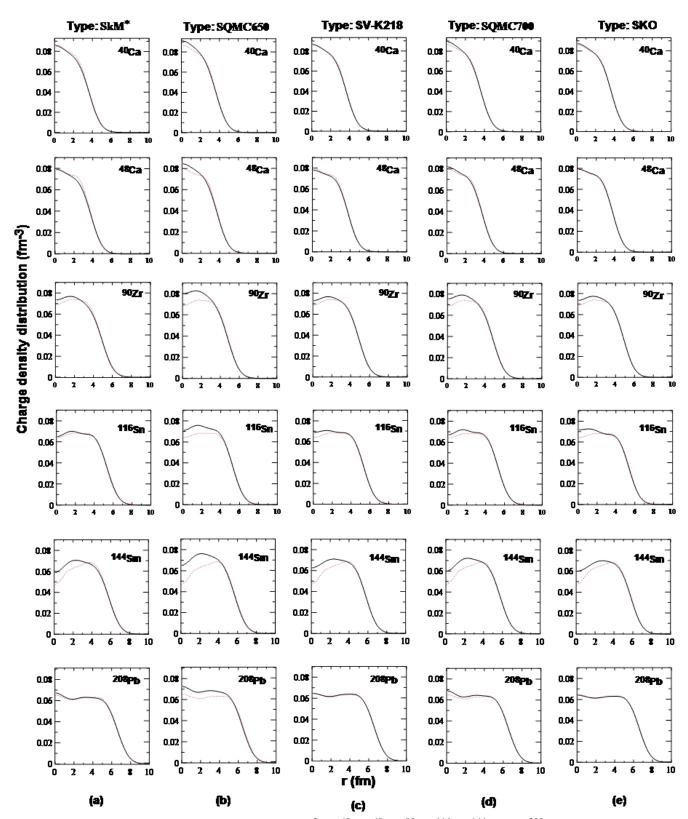
For <sup>116</sup>Sn, our results of charge distribution for all interactions were in a good agreement with the experimental data [8] in the center and at the surface of the nucleus but not so good inside the nucleus as in SQMC650, SKO,  $Sk\chi m$  parameters in less overestimated the experimental data line inside the nucleus and GS2, QMC2, SIII and SVII interactions underestimated the experimental data line inside the nucleus. The differences in describing data may due to differences in nuclear properties such as  $K_{NM}$ , where its value the first group is lower than of the second group.

For  ${}^{144}Sm$  our results of charge distribution for all interactions were in a good agreement with the experimental data [8] at the surface of the nucleus but not so good inside the nucleus.

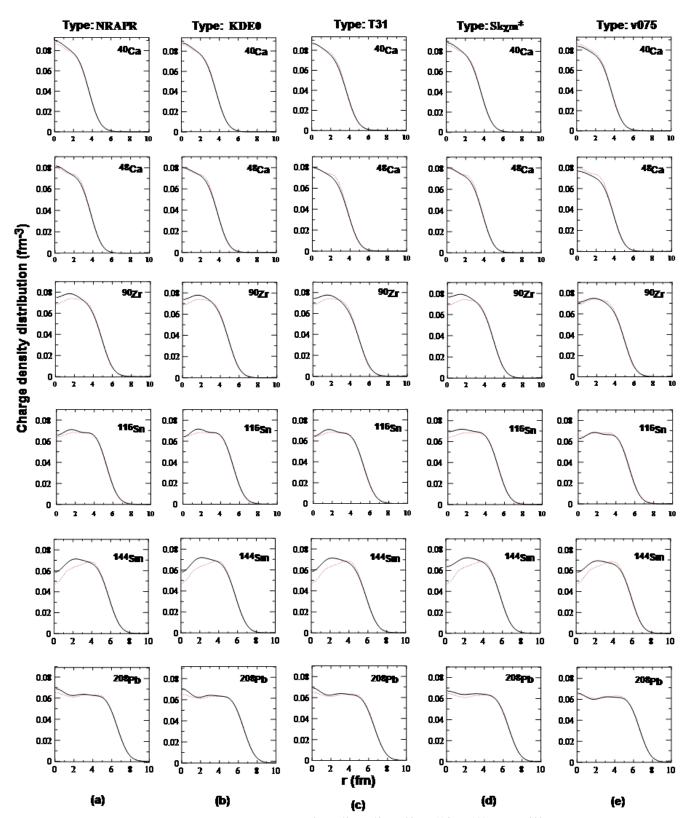
For <sup>208</sup>*Pb* our results of charge distribution for all interactions were in a good agreement with the experimental data [8] in the center and at the surface of the nucleus but not so good inside the nucleus as GS2, QMC2 and SVII and less for SIV and SIII interactions underestimated the experimental data line inside the nucleus and SQMC650, KDE0 parameters in less overestimated the experimental data line inside the nucleus. This may be due to differences in the nuclear matter properties of the investigated types of Skyrme interaction. By considering a light, medium and heavy nuclei, we can illustrate the impact of the effect of incompressibility modulus  $K_{NM}$  and symmetry energy density J on root-mean-square radii of neutron  $R_n$  and proton  $R_p$  for light, medium and heavy closed-shell nuclei <sup>40,48</sup>Ca, <sup>90</sup>Zr, <sup>116</sup>Sn, <sup>144</sup>Sm and <sup>208</sup>Pb.

In Fig.5 the incompressibility  $K_{NM}$  are shown as function of  $R_n$  and  $R_p$  for  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei respectively. the linear correlation coefficients of  $K_{NM}$  versus  $R_n$ , within the Skyrme forces were 0.395, 0.369, 0.512\*, 0.515\*, 0.605\*\* and 0.544\* for  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei, respectively, where it is small (weak) in  ${}^{40}Ca$  and  ${}^{48}Ca$ . The linear correlation coefficients of  $K_{NM}$  versus  $R_p$  within the Skyrme forces reaches 0.378, 0.489\*, 0.552\*, 0.562\*, 0.591\* and 0.606\*\* for  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei, respectively, also it is small (weak) in  ${}^{40}Ca$  and  ${}^{48}Ca$ . From Fig.5, we conclude there is an impact of the incompressibility  $K_{NM}$  on  $R_n$  and  $R_p$ .

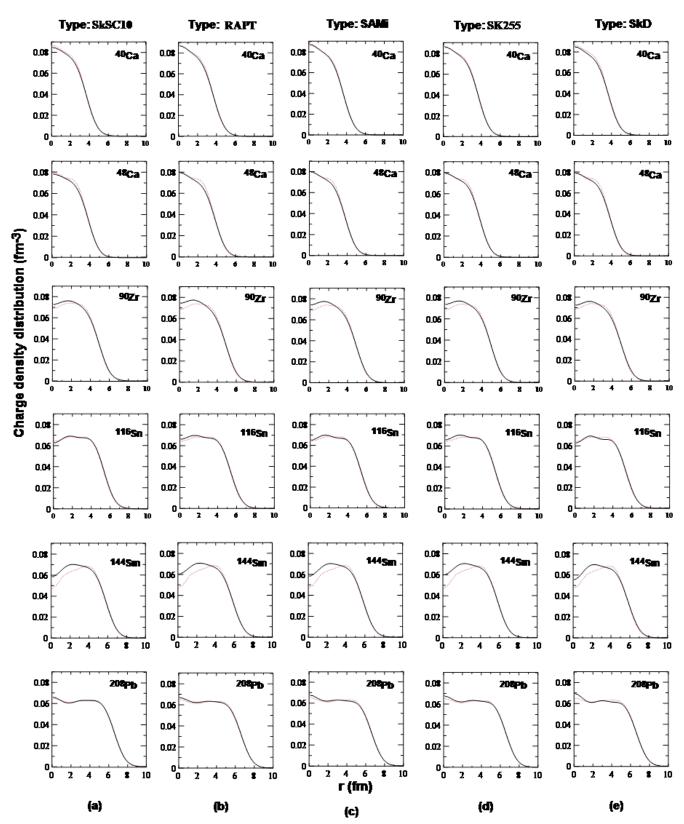
In Fig.6, the symmetry energy density J are shown as function of  $R_n$  and  $R_p$  for  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei respectively within the Skyrme forces,



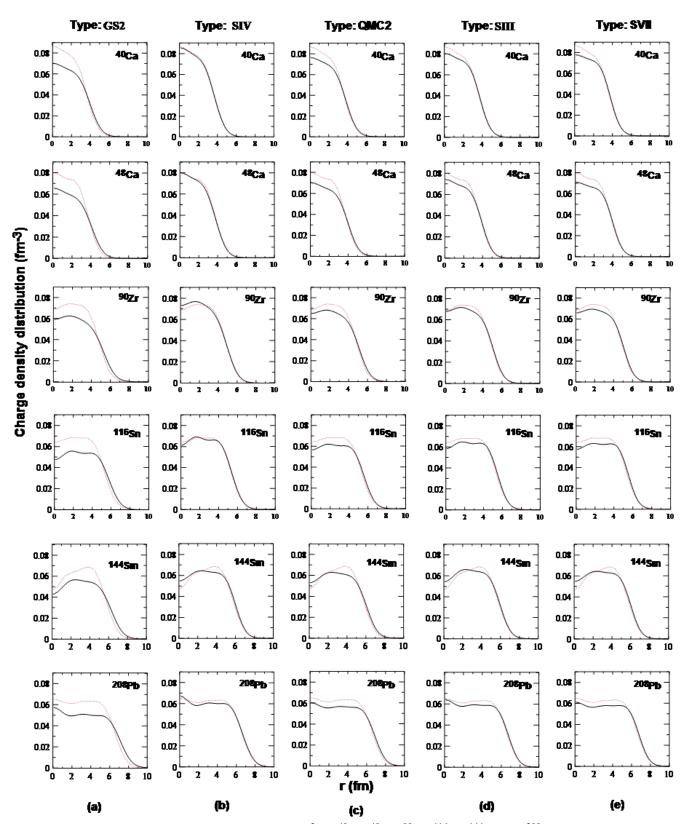
**Figure 1.** Our obtained charge density distribution  $(fm^{-3})$  of  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei with Skyrme interactions  $Skm^*$ , SQMC650, SV-K218, SQMC700, SKO. Charge densities are plotted by black solid line. The red dotted line shows experimental data taken from Ref [8].



**Figure 2.** Our obtained charge density distribution  $(fm^{-3})$  of  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei with Skyrme interactions NRAPER, KDE0, T31 *Sk* $\chi$ m and v075. Charge densities are plotted by black solid line. The red dotted line shows experimental data taken from Ref. [8].



**Figure 3.** Our obtained charge density distribution  $(fm^{-3})$  of  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei with Skyrme interactions SKSC10, RATP, SAMi, SK255 and SKD. Charge densities are plotted by black solid line. The red dotted line shows experimental data taken from Ref.[8].



**Figure 4.** Our obtained charge density distribution  $(fm^{-3})$  of  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei with Skyrme interactions GS2, SIV, QMC2, SIII and SVII. Charge densities are plotted by black solid line. The red dotted line shows experimental data taken from Ref.[8].

the linear correlation coefficients of J versus  $R_n$  were -0.433-, -0.179-, -0.308-, -0.254-, -0.212- and -0.032- for  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei, respectively and it is very small correlated except for  ${}^{40}Ca$  it is rather better than other. The linear correlation coefficients of J versus  $R_p$  were  $-0.462^*$ ,  $-0.497^*$ ,  $-0.496^*$ ,  $-0.499^*$ ,  $-0.475^*$  and  $-0.476^*$  for  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$ and <sup>208</sup>Pb nuclei, respectively are small (weak). From Fig.6, It can be concluded that the impact of the symmetry energy density J on neutron radii  $(J \rightarrow R_n)$  is weaker than that of  $R_n$  $(J \rightarrow R_p)$  or nonexistent. In general, from Figs.5 and 6, there is an influence of two nuclear matter properties (J and  $K_{NM}$ ) on proton and neutron radii in a space of Skyrme functional, we demonstrate the existence of a relation between  $(K_{NM})$  $\rightarrow$  R<sub>n</sub> and R<sub>p</sub>) for all the investigated nuclei except for <sup>40</sup>Ca and  ${}^{48}Ca$   $(K_{NM} \rightarrow R_n)$  it is weak. This illustrates that these quantities are coupled by the skyrme force and higher than  $(J \rightarrow R_n, R_p)$ , where the relation  $(J \rightarrow R_n)$  is smaller (weaker) than  $(J \rightarrow R_p)$ .

# 4. Conclusions:

In this study, the effect of incompressibility modulus  $K_{NM}$ and symmetry energy density J on charge distribution and root-mean-square radii of neutron  $R_n$  and proton  $R_p$  has been investigated for light, medium and heavy closed-shell nuclei  ${}^{40,48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  within the framework of self-consistent Hartree-Fock (HF) with 20 types of Skyrme interaction. Good agreements were obtained between our HF results of charge density distribution for all of the used types of Skyrme interaction with the experimental data at the surface and interior regions, while our results underestimated the experimental data line inside the investigated nuclei for SVII, SIII, QMC2 and GS2 sets. Concerning, the sensitivity of proton and neutron root-mean-square radii to nuclear matter properties, we conclude there is an impact of the incompressibility  $K_{NM}$  on  $R_n$  and  $R_p$ . Also, it can be concluded that the impact of the symmetry energy density J on neutron radii (J  $\rightarrow \mathbf{R}_n$ ) is weaker than that of  $R_p$  (J  $\rightarrow \mathbf{R}_p$ ) or nonexistent.

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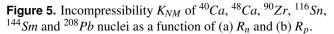
**Data Availability Statement:** All of the data supporting the findings of the presented study are available from corresponding author on request.

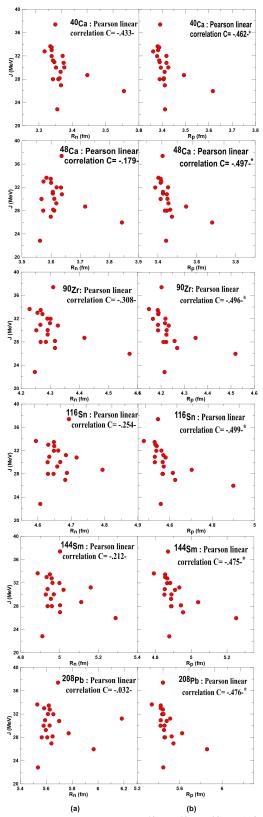
# **Declarations:**

**Conflict of interest:** The authors declare that they have no conflict of interest.

**Ethical approval:** The manuscript has not been published or submitted to another journal, nor is it under review.

40Ca : Pearson linear 40Ca : Pearson linear 360 correlation C= .395 correlation C=.378 () 320 Me ¥ 280 240 200 2 2 3.5 Rn (fm) 3.6 3.7 3.3 3.6 3.7 400 48Ca : Pearson linear 48Ca : Pearson linear 360 correlation C=.369 correlation C=.489 () 320 Me/ ¥ 280 240 200 – 3.5 3.8 3.6 Rp (fm) 3.8 Rn (fm) 400 90Zr: Pearson linear 90Zr: Pearson linear 360 correlation C= .512\* correlation C= .552' (MeV) MeV , INN 280 200 4.2 4.4 Rn (fm) 4.5 4.6 4.2 4.3 4.3 Rp (fm) 4.5 400 116Sn : Pearson linear 116Sn : Pearson linear correlation C= .562 correlation C= .515 360 WN 280 240 200 Rn (fm) Rp (fm) 400 144Sm : Pearson linear 144Sm : Pearson linear correlation C= .605\* correlation C= .591 360 32 MeV ¥ 280 240 200 Rn (fm) Rp (fm) 400 208Pb : Pearson linear 208Pb : Pearson linear correlation C= .544\* correlation C= .606 360 () 320 We <sup>WN</sup> 280 240 200 L 5.4 6.2 5.2 5.8 Rn (fm) 5.6 Rp (fm) (a) (b)





**Figure 6.** Symmetry energy J of  ${}^{40}Ca$ ,  ${}^{48}Ca$ ,  ${}^{90}Zr$ ,  ${}^{116}Sn$ ,  ${}^{144}Sm$  and  ${}^{208}Pb$  nuclei as a function of (a)  $R_n$  and (b)  $R_p$ .

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### الخلاصة

في العمل الحالي ، تم التحقق من تأثير معامل عدم الانضغاط *K<sub>NM</sub> وك*ثافة طاقة التناظر *L* على توزيع الشحنة ونصف قطر الجذر التربيعي للنيوترون *R<sub>n</sub> والبروتون R<sub>p</sub> لجموعة واسعة من انوية الذرات مغلقة الغلاف الخفيفة والمتوسطة والثقيلة 2020<sup>40</sup> ب 40<sup>20</sup>2<sub>20</sub> ، 20<sup>20</sup><sup>40</sup><sup>40</sup> ، 20<sup>40</sup><sup>40</sup><sup>50</sup> و 20<sup>40</sup><sup>40</sup><sup>40</sup> و 20<sup>40</sup><sup>40</sup><sup>40</sup><sup>50</sup> في إطار تقريب (Hartree – Fock(HF) المتسق ذاتيًا مع 20 نوعًا من تفاعلات <i>Kyrme بلو 10<sup>40</sup><sup>40</sup><sup>40</sup><sup>50</sup><sup>40</sup><sup>50</sup><sup>50</sup> و 20<sup>40</sup><sup>40</sup><sup>40</sup><sup>50</sup><sup>50</sup> في إطار تقريب (K<sub>M</sub>) عدم الانضغاط <i>K<sub>NM</sub> وك*ثافة طاقة التناظر *L* . تم تفاعلات *Kyrme المحموعة واسعة من الخصائص النووية مثل مع*امل عدم الانضغاط *K<sub>NM</sub> وك*ثافة طاقة التناظر *L* . تم الحصول على كثافات الشحنة النووية ومقارتها مع البيانات التجربة لكي تعطينا صورة عن البنية الداخلية للانوية التي تم فصها. كما تم دراسة العلاقة بين نصف قطر الجذر التربيعي للبروتون والنيترون مع خصائص المادة النووية والتي شملت معامل عدم الانضغاط *K<sub>NM</sub> وك*ثافة طاقة التناظر *L* من خلال ايجاد ومناقشة معامل ارتباط بيرسون الاحصائي الخطي لعرض الساعدة في الانضغاط المرابع النووي.

> **الكلمات الدالة:** تقريب (Hartree – Fock(HF ، تفاعل Skyrme ، توزيع الشحنة، انصاف اقطار الحذر التربيعي. **التمويل: لايوجد. بيان توفر البيانات: ج**ميع البيانات الداعمة لنتائج الدراسة المقدمة يمكن طلبها من المؤلف السؤول. **اقرارات: تضارب المصالح**: يقر المؤلفون أنه ليس لديهم تضارب في المصالح.

**الموافقة الأخلاقية:** لم يتم نشر المخطوطة أو تقديمها لمجلة أخرى، كما أنها ليست قيد المراجعة.